

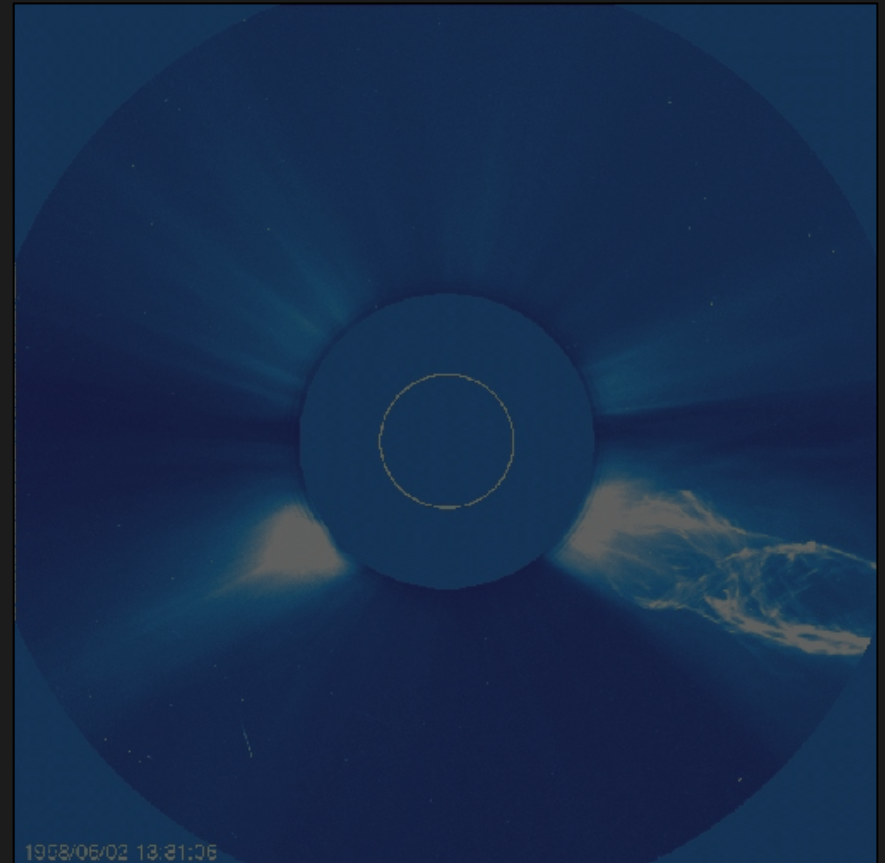
Magnetic helicity measurements in the solar wind

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G. Kleindienst, Ph. Bourdin,
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Helical CME, June 2, 1998
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Take-home messages

Magnetic helicity (density) ... $\mathbf{A} \cdot \mathbf{B}$

- * unambiguous observation as a challenge!
- * apparently random (in sign) in a typical solar wind (judging from time series data in the ecliptic plane)
- * plasma-physical effects interfering helicity transport (L-mode excitation, damping, polarization reversal)
- * Parker Solar Probe, Solar Orbiter, BepiColombo missions

Evaluating the helicity

Magnetic helicity

$$H^M = \int \mathbf{A} \cdot \mathbf{B} dV$$

Coulomb-gauge relation

$$\nabla \times \mathbf{B} = -\nabla^2 \mathbf{A}$$

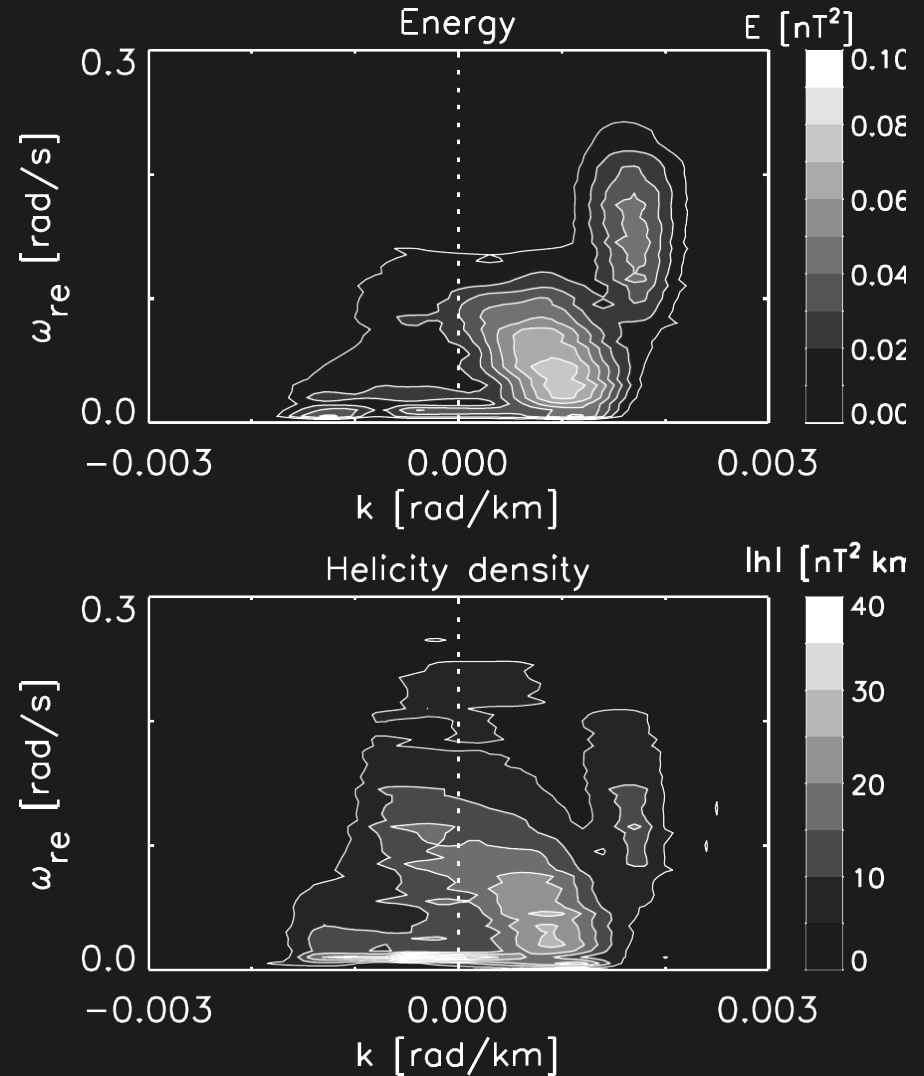
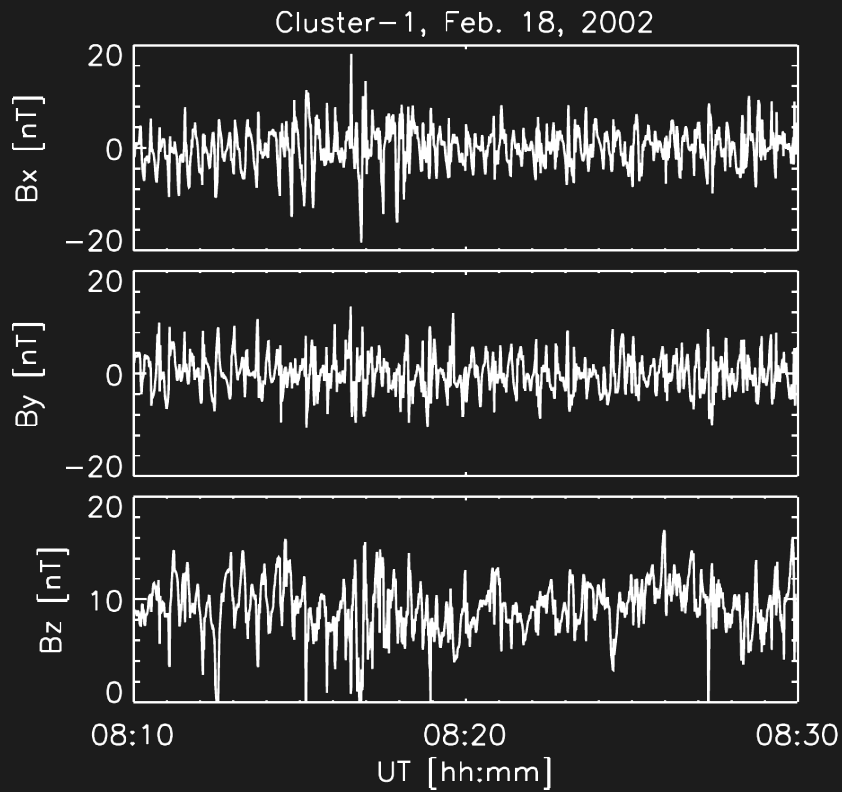
Un-curling in the Fourier space

$$\mathbf{a} = -\frac{i}{k^2} \mathbf{k} \times \mathbf{b}$$

Helicity density in the Fourier space

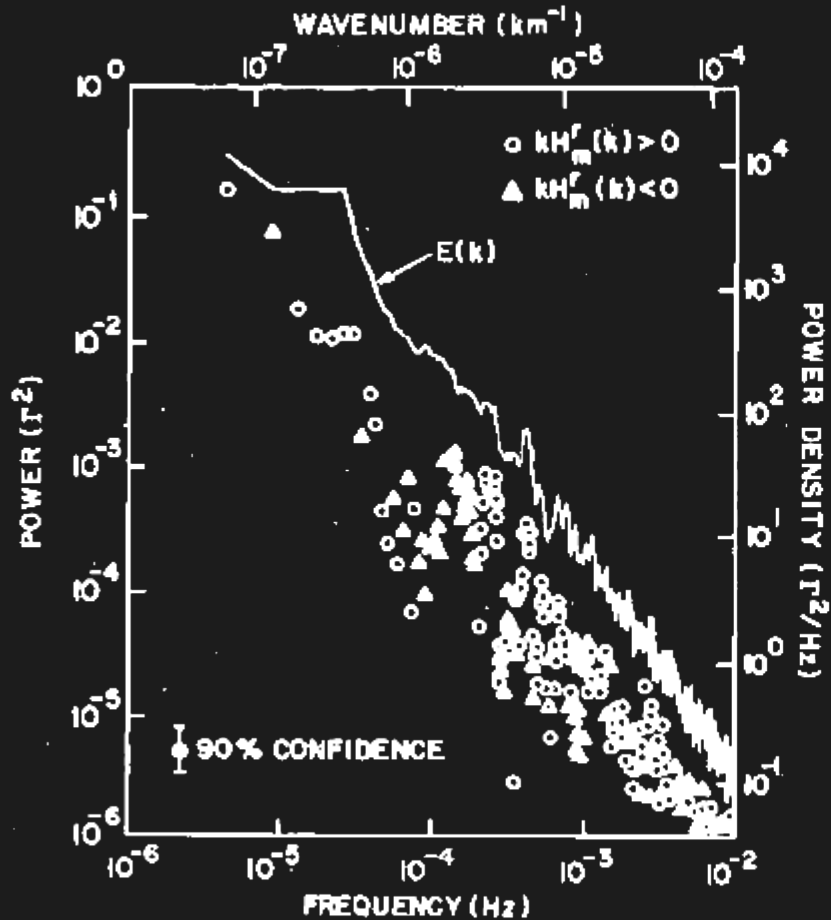
$$\langle \mathbf{a}^\dagger \cdot \mathbf{b} \rangle = -\frac{i}{k^2} \left[k_x \left(\langle b_y^* b_z \rangle - \langle b_z^* b_y \rangle \right) + k_y \left(\langle b_z^* b_x \rangle - \langle b_x^* b_z \rangle \right) \right. \\ \left. + k_z \left(\langle b_x^* b_y \rangle - \langle b_y^* b_x \rangle \right) \right].$$

Helicity using 4-point data

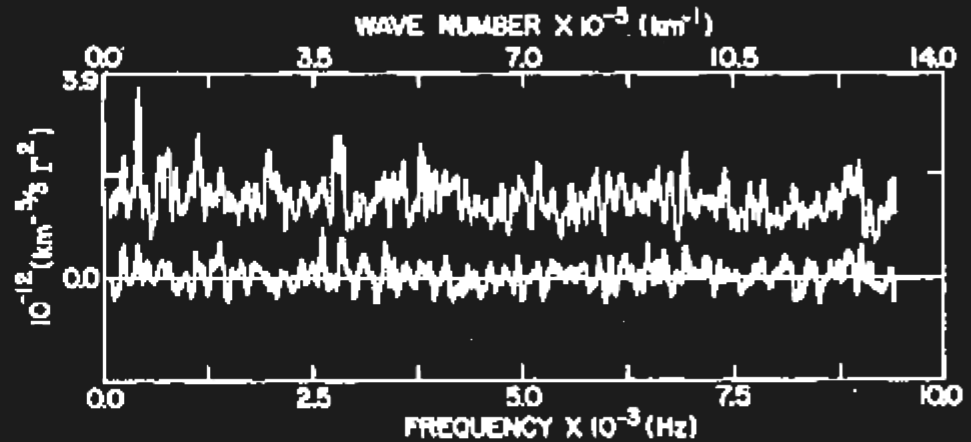


Narita et al., Ann. Geophys. 2009

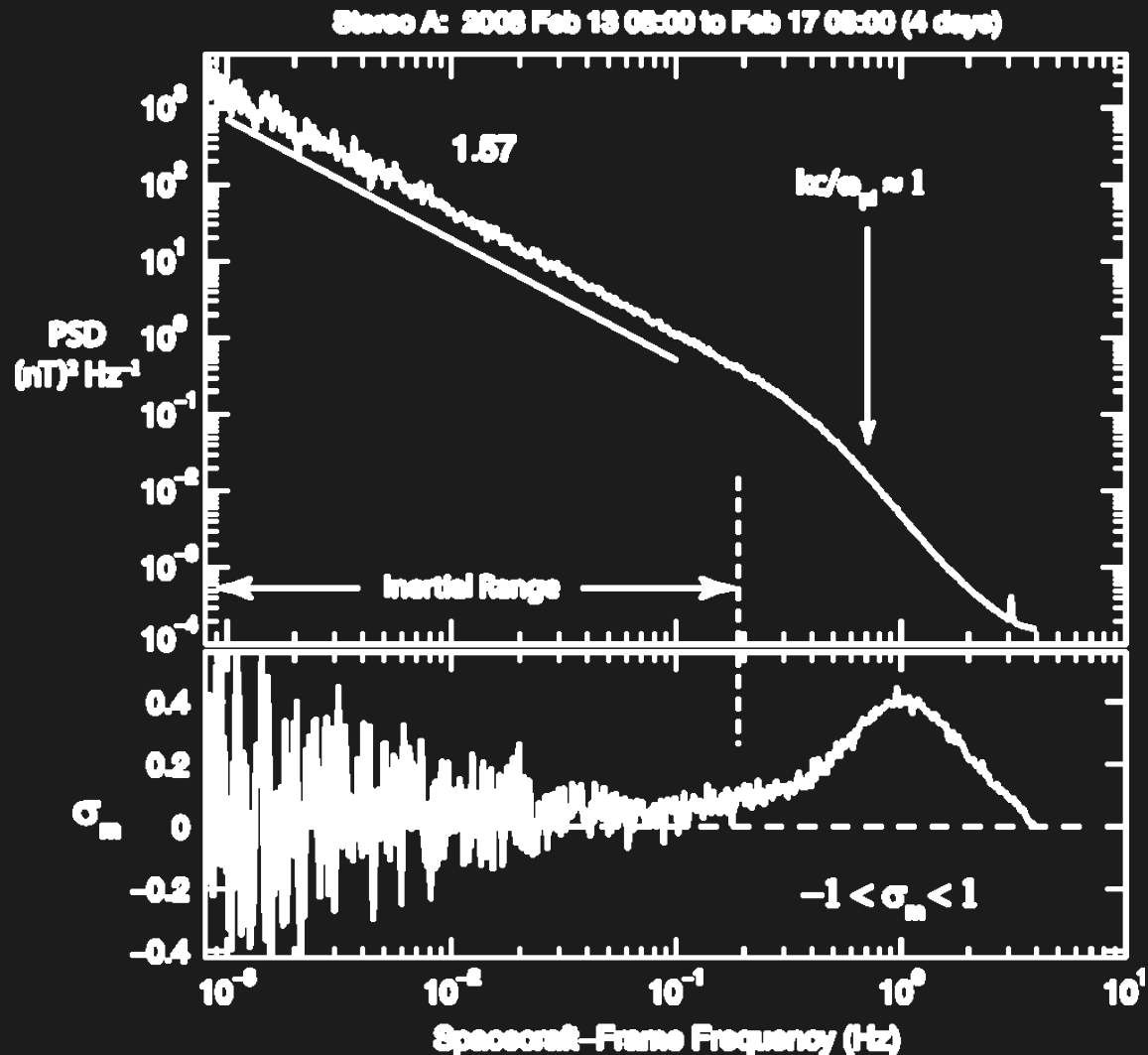
One-point data (more conventional)



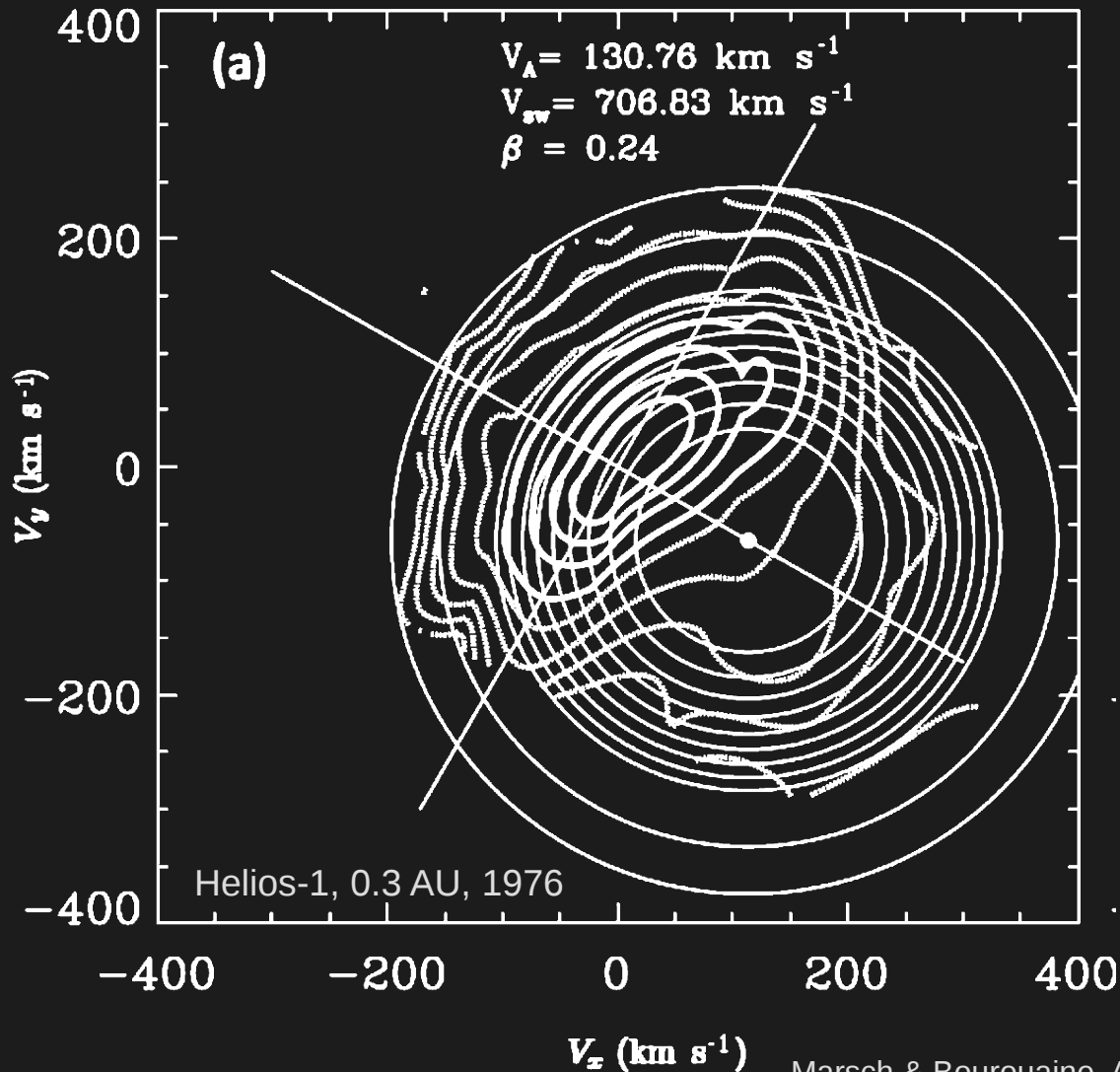
Voyager-2, 2.8AU, 1978



L-mode disappears on 100-km scale

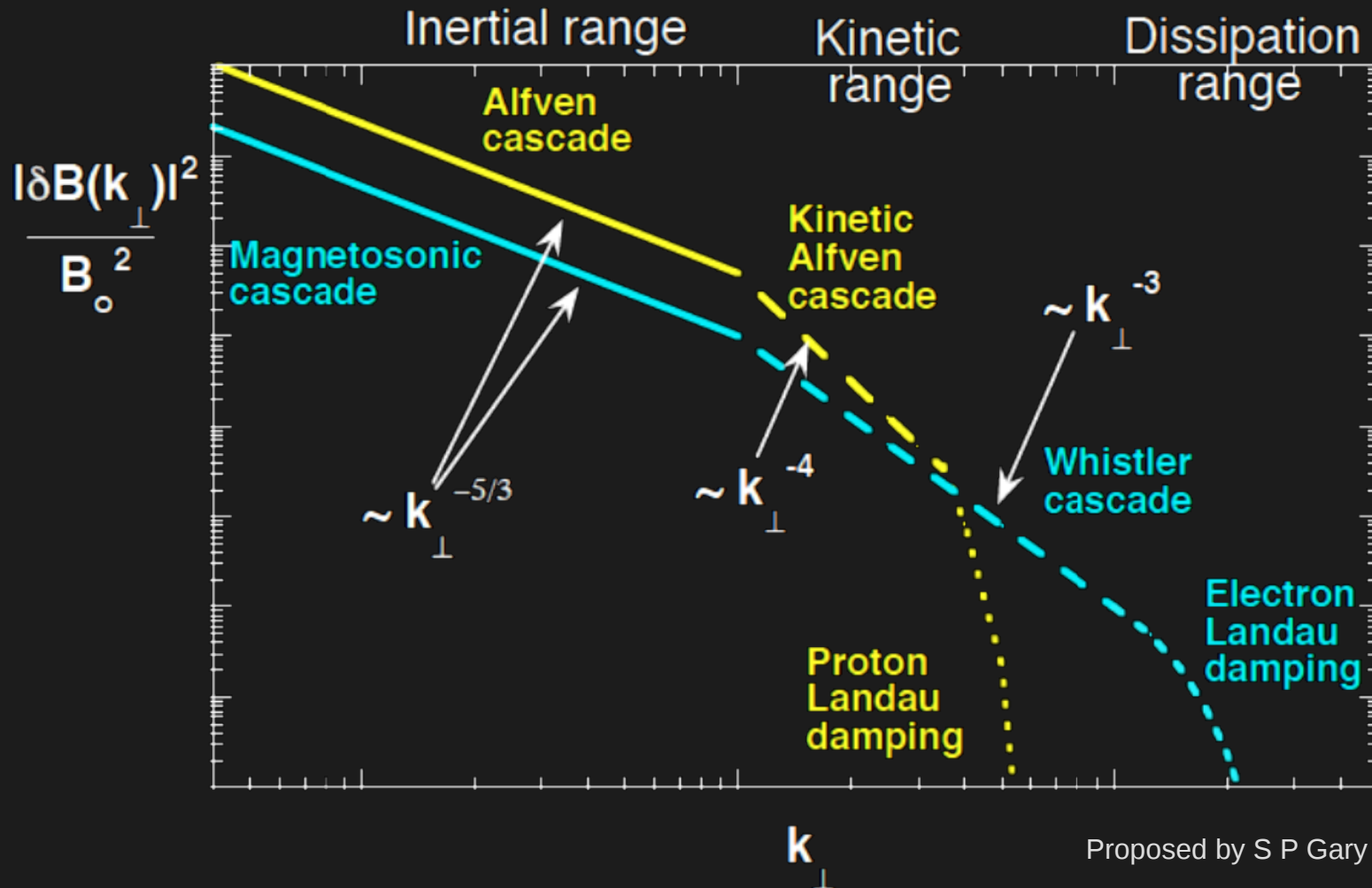


Pitch-angle scattering by L-mode waves



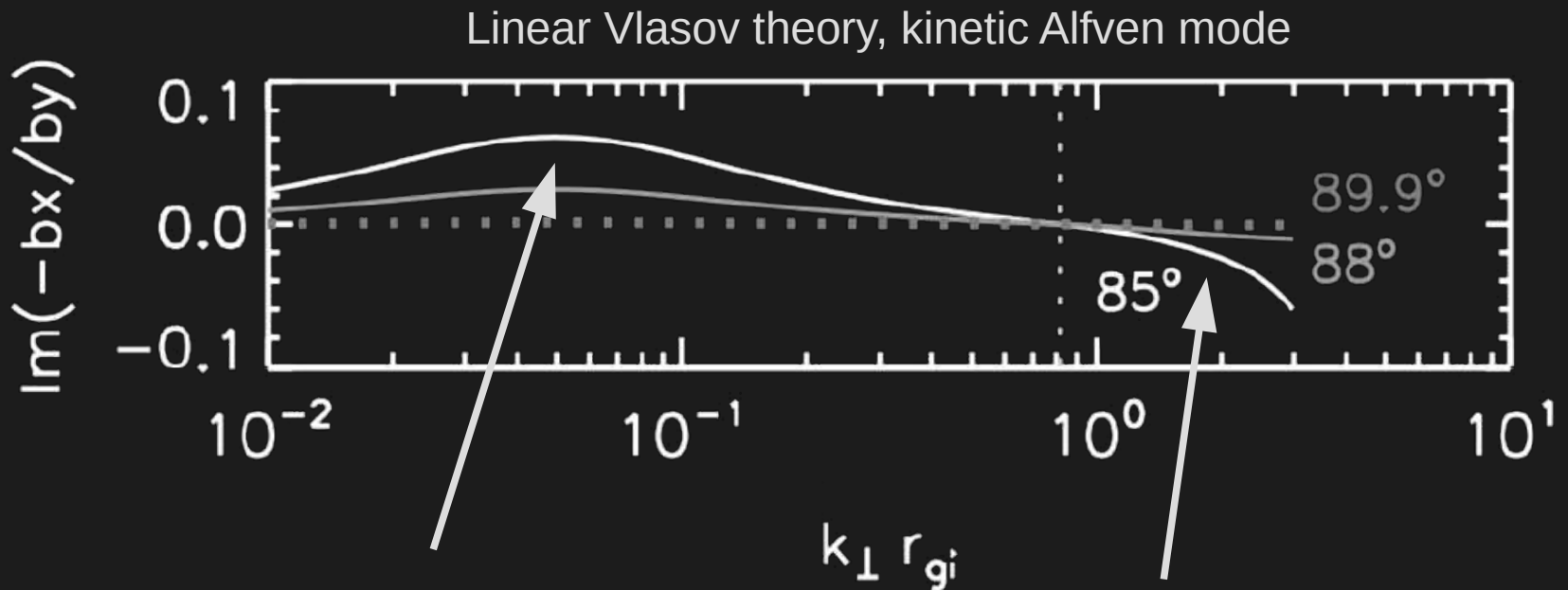
Marsch & Bourouaine, Ann. Geophys. 2011
Marsch, Ann. Geophys. 2018

A model of turb. energy spectrum



Proposed by S P Gary (2016)

Polarization reversal by linear mode



Hall current dominates

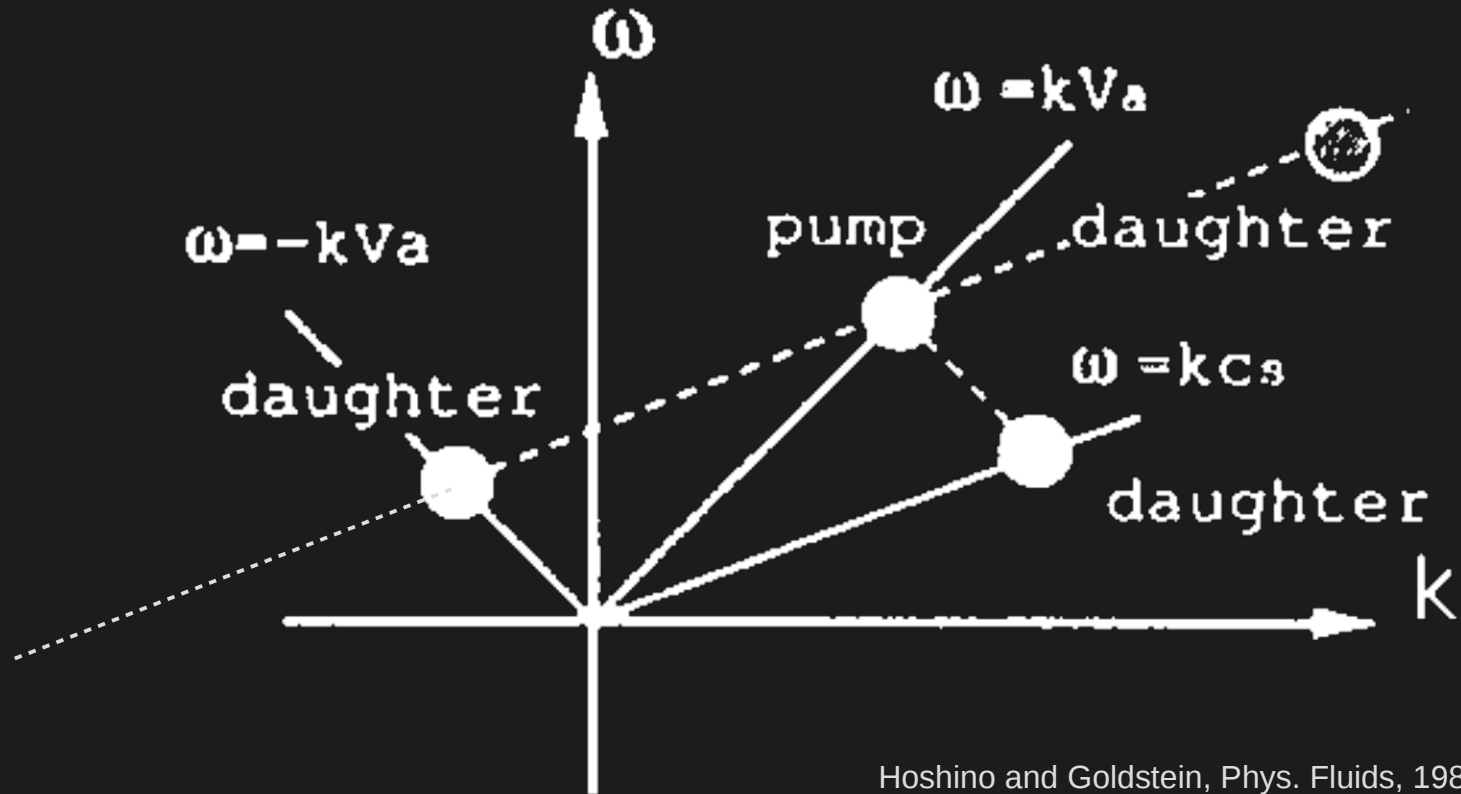
$$E_y = -\frac{1}{n_e e} j_x B_0$$

Diamagnetic current dominates

$$\vec{j} = \frac{\vec{B} \times \nabla p}{B^2}$$

Narita et al., Frontiers Phys. 2020

Polarization reversal by nonlinear mode



Hoshino and Goldstein, Phys. Fluids, 1989